

## Lecture 1. Basic concepts (book: p.1-19; 28-34)

Reference body. 'Observe' = deduce. Postulates. Standard configuration.

Spacetime diagrams; axes and simultaneity

Lorentz transformation:

$$\left. \begin{aligned} t' &= \gamma(t - vx/c^2) \\ x' &= \gamma(-vt + x) \\ y' &= y \\ z' &= z \end{aligned} \right\}, \quad \begin{pmatrix} ct' \\ x' \\ y' \\ z' \end{pmatrix} = \begin{pmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} ct \\ x \\ y \\ z \end{pmatrix}$$

Minkowski metric and definition of  $\Lambda$ :

$$g = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}, \quad \begin{aligned} \Lambda^T g \Lambda &= g \\ \Rightarrow \mathbf{A}'^T g \mathbf{B}' &= \mathbf{A}^T g \mathbf{B} \equiv \mathbf{A} \cdot \mathbf{B} \quad \text{Invariant} \end{aligned}$$

**Postulate 1, “Principle of Relativity”:**

The motions of bodies included in a given space are the same among themselves, whether that space is at rest or moves uniformly forward in a straight line.

**Postulate 2, “Light speed postulate”:**

There is a finite maximum speed for signals.

Alternative statements:

*Postulate 1:*

The laws of physics take the same mathematical form in all inertial frames of reference.

*Postulate 2:*

There is an inertial reference frame in which the speed of light in vacuum is independent of the motion of the source.

**Lecture 2. 4-vectors; Proper time; Method of invariants; transformation of velocity** (book p.14; 399-400; 22-25; 39-44)

Proper time

$$\frac{dt}{d\tau} = \gamma$$

Familiarity with  $\gamma$  :

$$\gamma = (1 - \beta^2)^{-1/2}, \quad \frac{d\gamma}{dv} = \gamma^3 v/c^2, \quad \frac{d}{dv}(\gamma v) = \gamma^3$$

symbol	definition	components	name(s)	invariant
X	X	$(ct, \mathbf{r})$	4-displacement, interval	$-c^2\tau^2$
U	$d\mathbf{X}/d\tau$	$(\gamma c, \gamma \mathbf{v})$	4-velocity	$-c^2$
P	$m_0 \mathbf{U}$	$(E/c, \mathbf{p})$	energy-momentum, 4-momentum	$-m_0^2 c^2$
A	$d\mathbf{U}/d\tau$	$\gamma(\dot{\gamma}c, \dot{\gamma}\mathbf{v} + \gamma\mathbf{a})$	4-acceleration	$a_0^2$

Timelike  $\mathbf{U} \cdot \mathbf{U} < 0$  e.g. 4-velocity

spacelike  $\mathbf{A} \cdot \mathbf{A} > 0$  e.g. 4-acceleration

null  $\mathbf{P} \cdot \mathbf{P} = 0$  e.g. energy-momentum of light pulse

**Method of invariants** = “Try using an invariant if you can, and pick an easy reference frame.”

4-acceleration is orthogonal to 4-velocity:  $\mathbf{U} \cdot \mathbf{A} = 0$ .

Transformation of velocity:

$$\mathbf{u}'_{\parallel} = \frac{\mathbf{u}_{\parallel} - \mathbf{v}}{1 - \mathbf{u} \cdot \mathbf{v}/c^2}, \quad \mathbf{u}'_{\perp} = \frac{\mathbf{u}_{\perp}}{\gamma_v (1 - \mathbf{u} \cdot \mathbf{v}/c^2)}. \quad \underbrace{\hspace{1cm}} \quad \underbrace{\hspace{1cm}} \quad \longrightarrow \text{particle jets}$$



### Lecture 3. Rapidity; Doppler effect; Headlight effect

1. **Rapidity:**  $\tanh(\rho) \equiv \frac{v}{c}, \Rightarrow \cosh(\rho) = \gamma, \quad \sinh(\rho) = \beta\gamma, \quad e^\rho = \left(\frac{1+\beta}{1-\beta}\right)^{1/2}.$

$$\Lambda = \begin{pmatrix} \cosh \rho & -\sinh \rho & 0 & 0 \\ -\sinh \rho & \cosh \rho & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}, \quad \begin{matrix} \boxed{1} & \boxed{2} \longrightarrow v_{21} & \boxed{3} \longrightarrow v_{32} \end{matrix}$$

Colinear velocities:  $v_{31} = \frac{v_{32} + v_{21}}{1 + v_{32}v_{21}/c^2}$

Colinear rapidities:  $\rho_{31} = \rho_{21} + \rho_{32}$

### 2. Doppler effect

4-wave vector  $\mathbf{K} \equiv (\omega/c, \mathbf{k}), \quad e^{i(\mathbf{k}\cdot\mathbf{r}-\omega t)} = e^{i\mathbf{K}\cdot\mathbf{X}}$

Doppler effect:  $\mathbf{K} \cdot \mathbf{U} \Rightarrow \frac{\omega}{\omega_0} = \frac{1}{\gamma(1 - (v/v_p) \cos \theta)}.$

### 3. Headlight effect or 'aberration':

Headlight effect:  $\mathbf{K} = \Lambda^{-1}\mathbf{K}_0 \Rightarrow \cos \theta = \frac{\cos \theta_0 + v/c}{1 + (v/c) \cos \theta_0} \Rightarrow \frac{d\Omega}{d\Omega_0} = \left(\frac{\omega_0}{\omega}\right)^2.$

→ **brightness** (power per unit solid angle) transforms as  $(\omega/\omega_0)^4$  for isotropic source

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## Lecture 4. Force; simple dynamical problems

1. Force

$$\mathbf{F} \equiv \frac{d\mathbf{P}}{d\tau}, \quad \mathbf{U} \cdot \mathbf{F} = \gamma^2 \left( -\frac{dE}{dt} + \mathbf{u} \cdot \mathbf{f} \right) = -c^2 \frac{dm_0}{d\tau}.$$

'Pure' force:

$$\mathbf{U} \cdot \mathbf{F} = 0 \Rightarrow m_0 = \text{const}, \quad \frac{dE}{dt} = \mathbf{f} \cdot \mathbf{u}$$

2. Transformation of force: use  $\Lambda\mathbf{F}$  and  $\gamma$ , or  $(d/dt')(\Lambda\mathbf{P})$  :

$$\mathbf{f}'_{\parallel} = \frac{\mathbf{f}_{\parallel} - (\mathbf{v}/c^2)dE/dt}{1 - \mathbf{u} \cdot \mathbf{v}/c^2}, \quad \mathbf{f}'_{\perp} = \frac{\mathbf{f}_{\perp}}{\gamma_v(1 - \mathbf{u} \cdot \mathbf{v}/c^2)} \Rightarrow \begin{cases} \mathbf{f}'_{\parallel} = \mathbf{f}_{\parallel} \\ \mathbf{f}'_{\perp} = \mathbf{f}_{\perp}/\gamma \end{cases} \text{ for } u = 0$$

3. Equation of motion in any given reference frame:

$$\mathbf{f} = \frac{d\mathbf{p}}{dt} = \frac{d}{dt}(\gamma m \mathbf{v}) = \gamma m \mathbf{a} + m \frac{d\gamma}{dt} \mathbf{v}$$

$$\Rightarrow \begin{cases} \text{(i)} & \text{acceleration is not necessarily parallel to force!} \\ \text{(ii)} & \mathbf{f} = m\mathbf{a} \text{ only valid at } \mathbf{v} = 0 \\ \text{(iii)} & \text{for 'pure' force } (m=\text{constant}): \quad f_{\parallel} = \gamma^3 m a_{\parallel}, \quad f_{\perp} = \gamma m a_{\perp}. \end{cases}$$

4. Uniform B field: just like Newtonian result, but with  $m$  replaced by  $\gamma m$ .

Hence  $\omega = qB/\gamma m$  and  $p = qBr$ .

5. Motion parallel to E field: hyperbolic motion,  $x^2 - c^2 t^2 = (c^2/a_0)^2$  and  $\ddot{\mathbf{U}} \propto \mathbf{U}$ .  
Constant proper acceleration.

## Lecture 5. Some more kinematics; The conservation of energy and momentum

1. Constant proper acceleration and hyperbolic motion:  $\frac{d\rho}{d\tau} = a_0$ .
  2. Rigidity; the Great Train Disaster
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3. Lewis and Tolman argument:  $\mathbf{p} = m\alpha(v)\mathbf{v} \Rightarrow \mathbf{p} = \gamma m\mathbf{v}$ .
4. Impact of simultaneity on “ $\mathbf{P}_{\text{tot}} = \mathbf{P}_1 + \mathbf{P}_2 + \mathbf{P}_3 + \dots + \mathbf{P}_n$ ”.
5. “Zero component lemma”: if one component of a 4-vector is zero in all reference frames, then the whole 4-vector is zero.

Hence momentum conservation  $\Leftrightarrow$  energy conservation.

Main postulates, momentum conservation  $\left. \vphantom{\begin{array}{l} \text{Main postulates,} \\ \text{momentum conservation} \end{array}} \right\} \Rightarrow E_0 = mc^2, \text{ equivalence of rest mass and rest energy.}$

## Lecture 6. Collisions.

Methods:

$$\mathbf{P} \cdot \mathbf{P} = -m^2 c^2 \Rightarrow E^2 - p^2 c^2 = m^2 c^4 \quad (1)$$

$$\mathbf{p} = \gamma m \mathbf{v}, \quad E = \gamma m c^2 \Rightarrow \mathbf{v} = \frac{\mathbf{p} c^2}{E} \quad (2)$$

$$\text{“Isolate and square.”} \quad (3)$$

1. Decay at rest. e.g. find the energy of one of the products:

$$E_1 = \frac{M^2 + m_1^2 - m_2^2}{2M} c^2, \quad E_{\text{photon}} = \Delta E_{\text{rest}} - E_{\text{recoil}}.$$

2. In-flight decay. e.g. Find the rest mass of the original particle:

$$M^2 = m_1^2 + m_2^2 + \frac{2}{c^4} (E_1 E_2 - \mathbf{p}_1 \cdot \mathbf{p}_2 c^2)$$

$\Rightarrow$  it suffices to measure  $m_1, m_2, p_1, p_2, \theta$ .

3. Particle formation. e.g. threshold energy (stationary target):

$$E_{\text{th}} = \frac{(\sum_i m_i)^2 - m^2 - M^2}{2M} c^2.$$

4. Elastic collision. e.g. find angles in the lab frame

5. Compton effect (A. H. Compton, Physical Review **21**, 483 (1923)):

$$\lambda' - \lambda = \frac{h}{mc} (1 - \cos \theta).$$

[other: 3-body decay, inverse Compton effect, etc.]



## Lecture 7. Composite body; 4-gradient; flow

1. The idea of a composite body with a net energy and momentum, hence rest mass and velocity.  $\mathbf{p} = \sum_i \mathbf{p}_i$ ,  $E = \sum_i E_i$ ,  $\mathbf{P} \equiv (E/c, \mathbf{p})$ ,  $m \equiv \sqrt{-\mathbf{P}^2/c^2}$ ,  $\mathbf{v} = \mathbf{p}c^2/E$ .

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2. Concept of a Lorentz-invariant scalar field.

e.g.  $\mathbf{B} \cdot \mathbf{E}$ ,  $E^2 - c^2 B^2$ , but NOT charge density or potential energy.

3. 4-gradient operator:

$$\square \equiv \left( -\frac{1}{c} \frac{\partial}{\partial t}, \frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z} \right)$$

If  $\phi$  is a Lorentz-invariant scalar quantity, then:  $\square' \phi = \Lambda \square \phi$

**i.e.  $\square \phi$  is a 4-vector.**

4. 4-divergence

$$\square \cdot \mathbf{F} = \frac{1}{c} \frac{\partial F^0}{\partial t} + \nabla \cdot \mathbf{f}, \quad \square^2 = \square \cdot \square = -\frac{1}{c^2} \frac{\partial^2}{\partial t^2} + \nabla^2.$$

5. Wave phase  $\phi$  is Lorentz-invariant.

$$\begin{aligned} \Rightarrow \quad \mathbf{K} &\equiv \square \phi \text{ is a 4-vector} \\ \square^2 \phi &= 0 \text{ is the wave equation.} \end{aligned}$$

6. Flow and conservation:

$$\begin{array}{lll} \text{4-current density} & \mathbf{J} \equiv \rho_0 \mathbf{U} & = (\rho c, \mathbf{j}) \\ \text{continuity equation} & \square \cdot \mathbf{J} = 0. & \end{array}$$