

FMOS & LOFAR?

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LOFAR



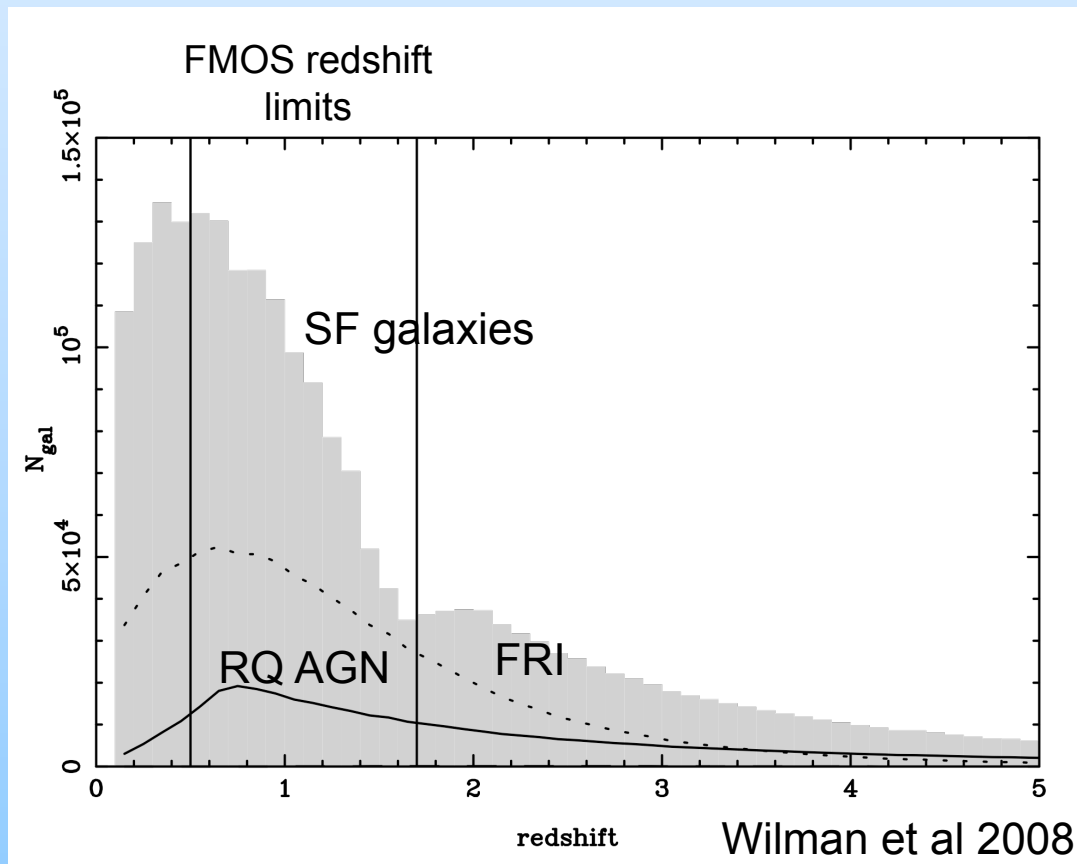
- LOFAR will be the fastest survey telescope in the world.
- Able to detect Milky Way-type galaxies up to $z \sim 3$, and SCUBA-type galaxies at $z > 6$
- Now being built in Netherlands.
- UK joined along with Germany.
- 1st in the new generation of powerful radio telescopes.
- Operates at 30-80MHz and 120-240MHz
- Free of any dust obscuration

LOFAR surveys

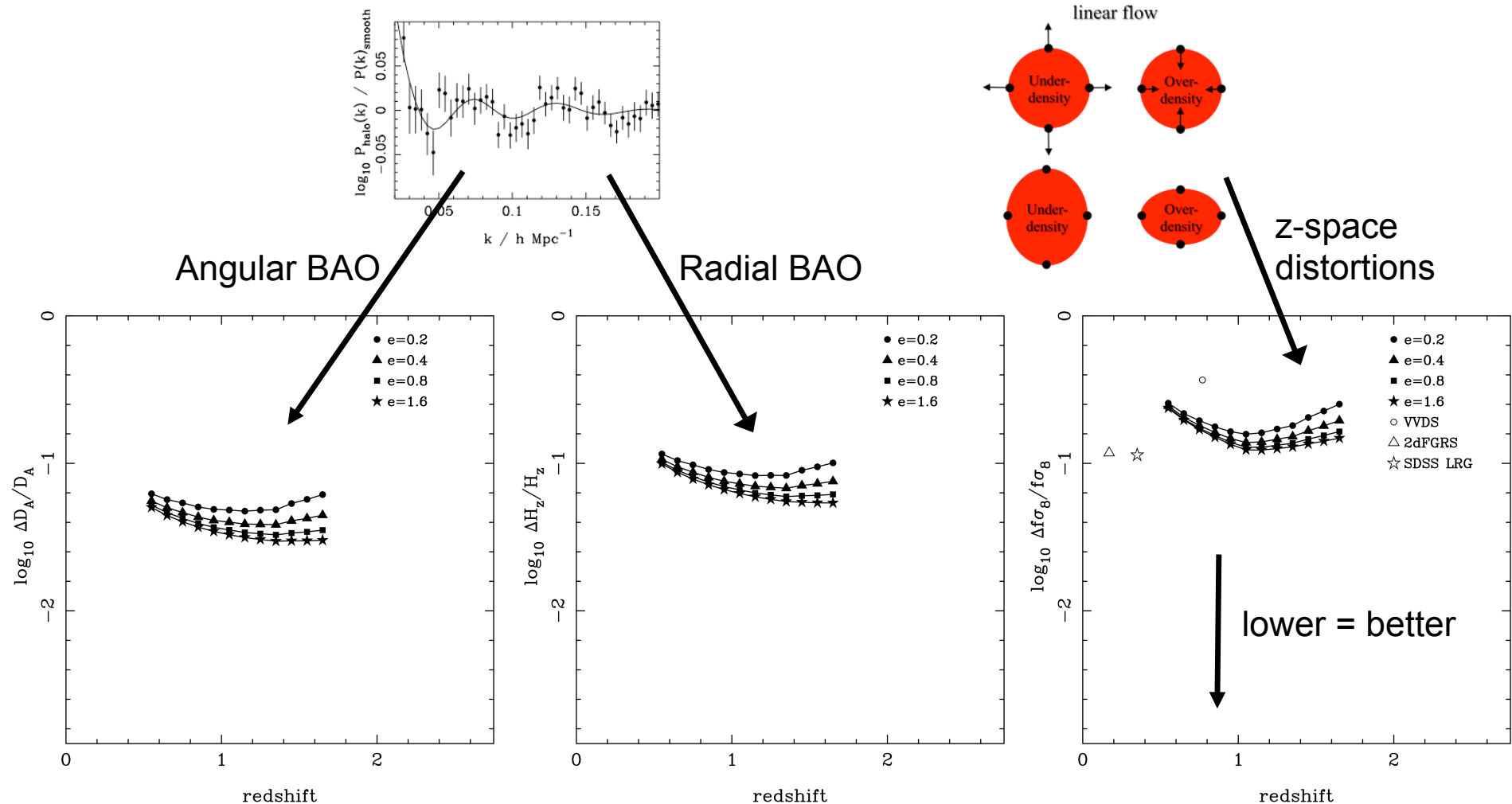
- All Sky Survey
 - 20,000 sq.degree survey at 15, 30, 60, 120, 200MHz to 10, 2, 0.75, 0.1, 0.2mJy
 - 1000 sq.degree survey at 200MHz to 0.065mJy (Cluster relics/haloes, starburst galaxies)
- Deep Survey
 - 3000 sq.deg at 30 & 60MHz to 0.7 & 0.25mJy
 - 550 sq.deg at 120MHz to 0.025mJy
 - 360 sq.deg.at 200MHz to 0.016mJy (distant starbursts, AGN, clusters...)
 - choose blank field regions with the best degree-scale multi-wavelength data
- Ultra-Deep Survey
 - 71 sq. deg. at 150MHz to 0.0062mJy (confusion limited at sub-arcsec resolution) very high-z starbursts, RQ-AGN, ...

LOFAR deep redshift distribution

- LOFAR 10σ sources expected in deep survey (550deg^2)
- cut to 300deg^2 survey
- assume we can remove $z < 0.4$ galaxies with photometric (SDSS?) selection
- gives $\sim 1 \times 10^6$ SF galaxies with $0.5 < z < 1.7$
- 1.9×10^6 other sources (including $0.4 < z < 0.5$ SF gals)
- gives a “redshift completeness” of 0.36
- pessimistic as would also get z for some AGN + extra photo- z selection. Also can use luminosity to favor SF galaxies



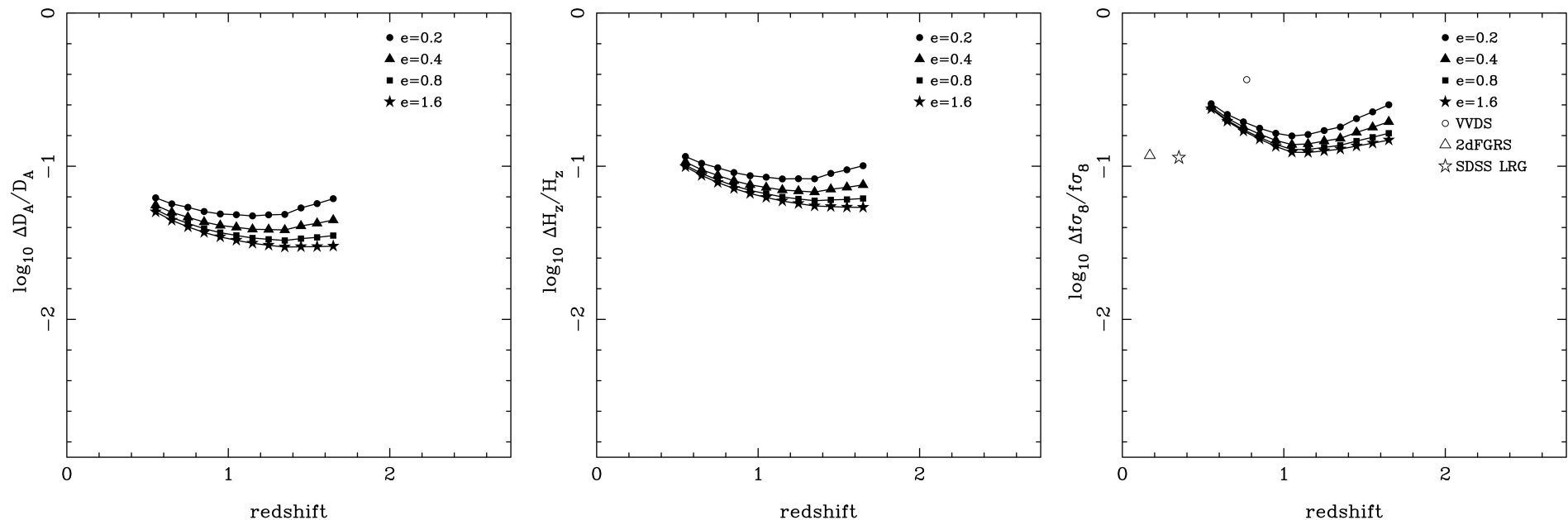
Plot explanation: cosmological constraints



- Fisher matrix predictions for cosmological constraints from BAO and from redshift-space distortions

Redshift completeness

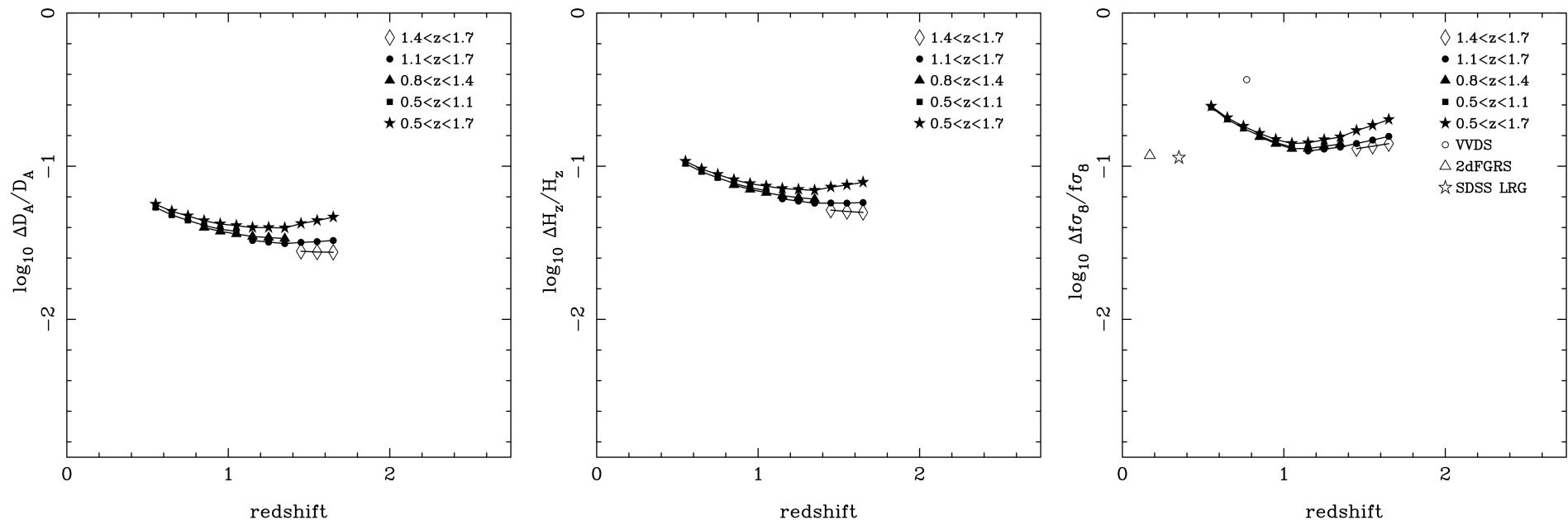
- take SF galaxy distribution (predicted 10σ LOFAR sources)
- $0.5 < z < 1.7$, 300deg² baseline FMOS survey has $\sim 600\,000$ targets
- effect of sub-sampling by a factor e shown below



- big gains until $e \sim 0.5$, then diminishing returns, particularly for redshift-space distortions

Galaxy selection?

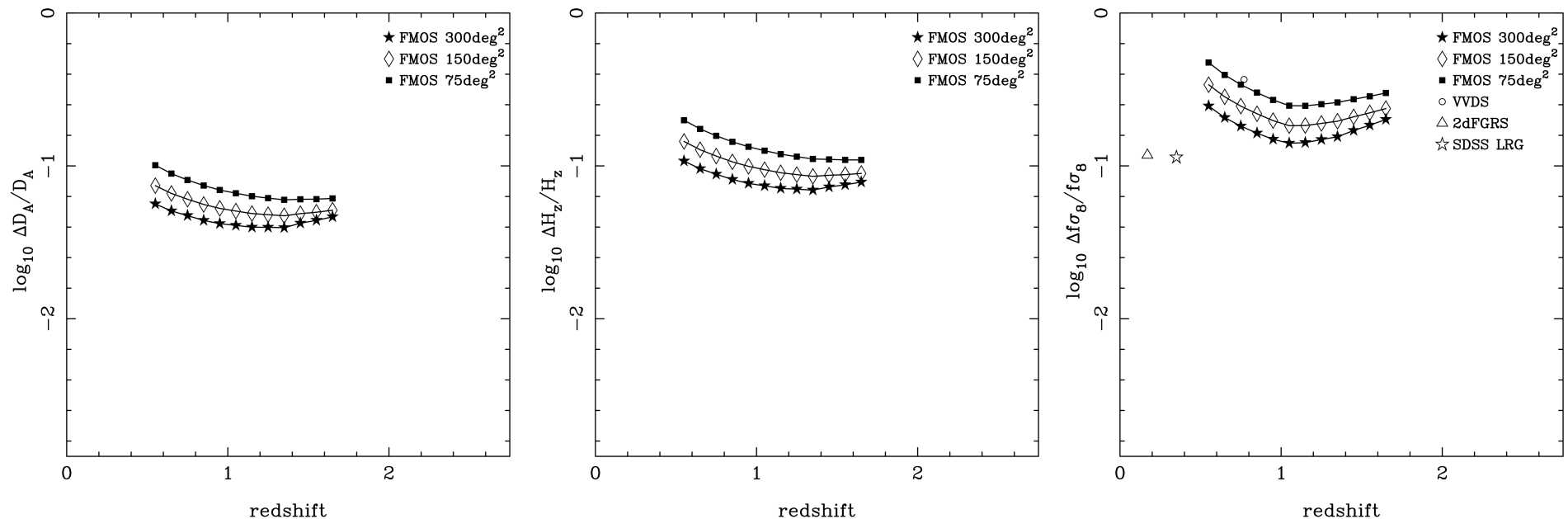
- take SF galaxy distribution (predicted 10σ LOFAR sources)
- $0.5 < z < 1.7$, 300deg^2 baseline FMOS survey has $\sim 600\,000$ targets
- assume fiducial $e=0.36$
- where to pre-select galaxies – ie put these galaxies in different redshift bins



- high number of high redshift galaxies helps

Volume vs number density

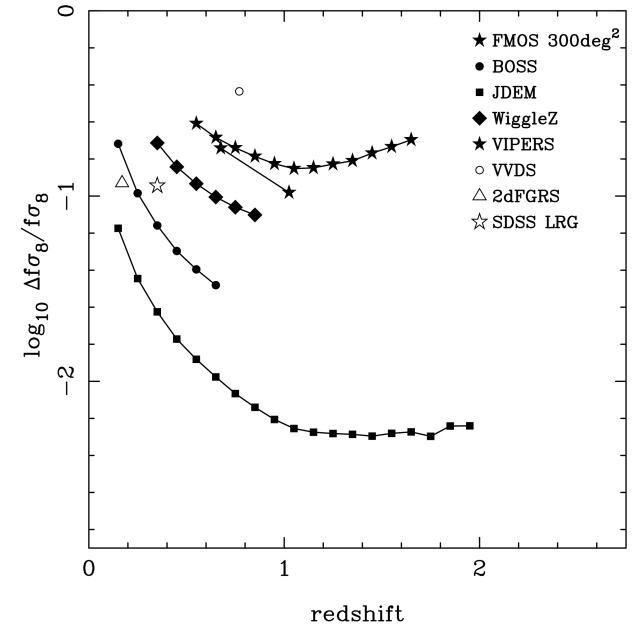
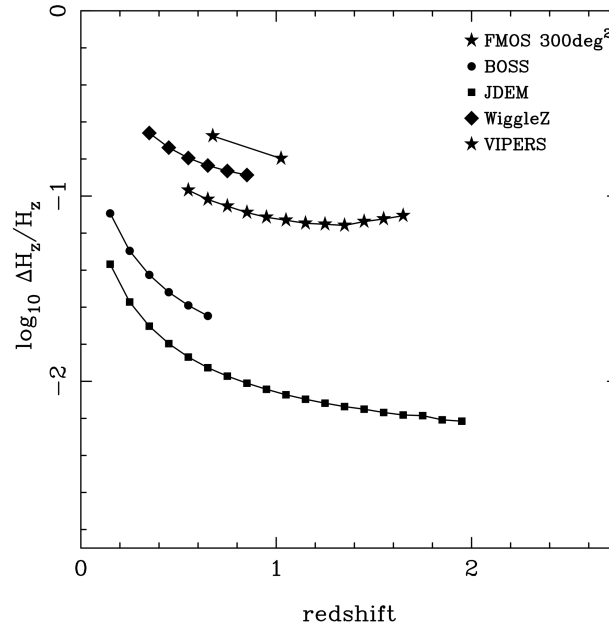
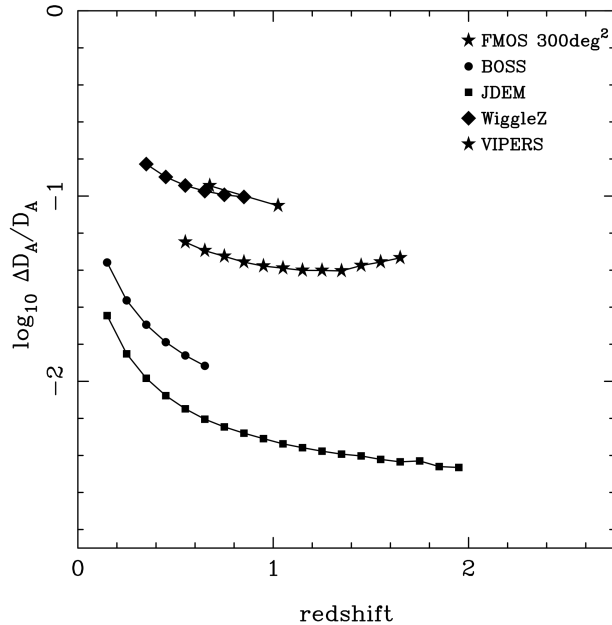
- take SF galaxy distribution (predicted 10σ LOFAR sources)
- $0.5 < z < 1.7$, 300deg² baseline FMOS survey has ~600 000 targets
- compare with oversampling by a factor of 2 or 4, covering a smaller area



- definitely do not want to decrease survey area

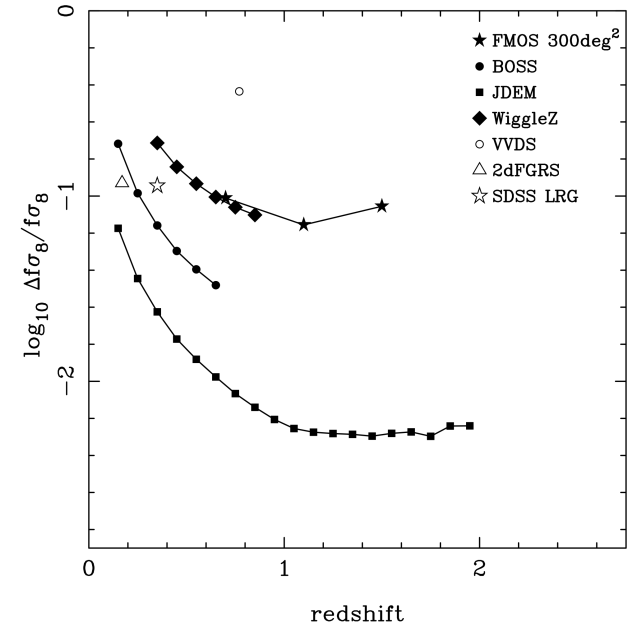
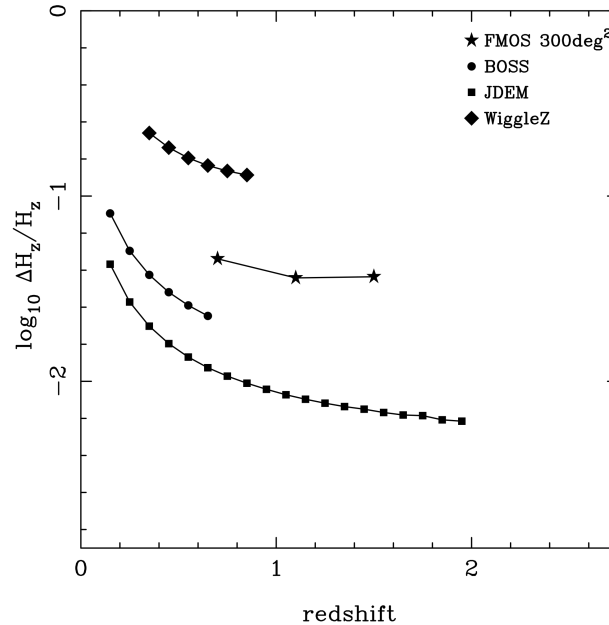
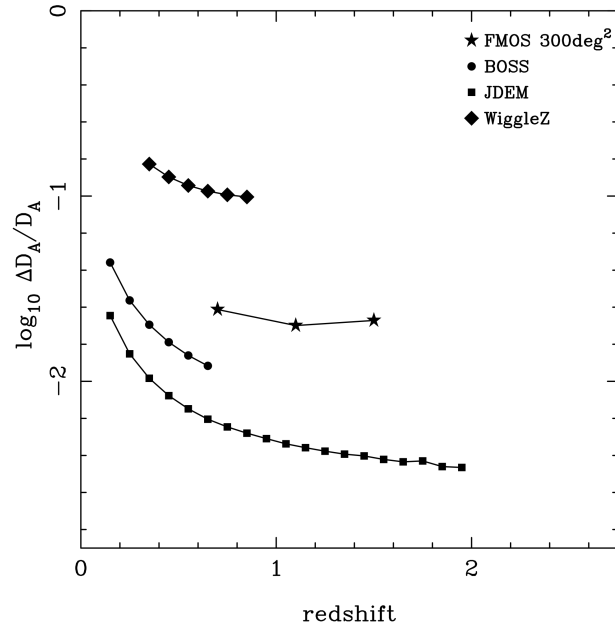
Comparison with other surveys

- take SF galaxy distribution (predicted 10σ LOFAR sources)
- $0.5 < z < 1.7$, 300deg^2 baseline FMOS survey has $\sim 600\,000$ targets
- assume can remove $z < 0.4$ galaxies from photometric redshifts
- leaves $e=0.36$ SF galaxy fraction in $0 < z < 1.7$
- pessimistic as can get redshifts for some AGN



Things look better with larger bins ...

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conclusions

- 300deg² FMOS survey means close to cosmic variance limit
- To optimise science return need to increase high-z galaxy distribution
- LOFAR deep selection can provide a sample of star-forming galaxies with a sampling return of 0.36, just removing $z < 0.4$ galaxies
- would also pick up redshifts for some of the AGN so $e=0.5-0.6$ probably more realistic